



Mark Hindmarsh (Sussex, Helsinki), Kari Rummukainen (Helsinki) and David J. Weir (Stavanger)

What was known

- ▶ Spontaneous breaking of a non-abelian symmetry can produce 'beads' consisting of 't Hooft-Polyakov monopoles, on cosmic strings.
- ▶ It is not known how the monopoles influence the dynamics of the resulting string network.

What this work adds

- ▶ We have carried out simulations of this scenario for the first time.
- ▶ Monopoles are carried along by the strings; the network behaves similarly to an abelian string network.

Next steps

- ▶ Study larger ratios between the monopole and string scales, to check whether monopoles eventually slow strings down.
- ▶ Observational predictions for strings in grand unified models, e.g. $SO(10)$.

Introduction

- ▶ We study the formation of cosmic string networks in the model with Lagrangian^{1,2}

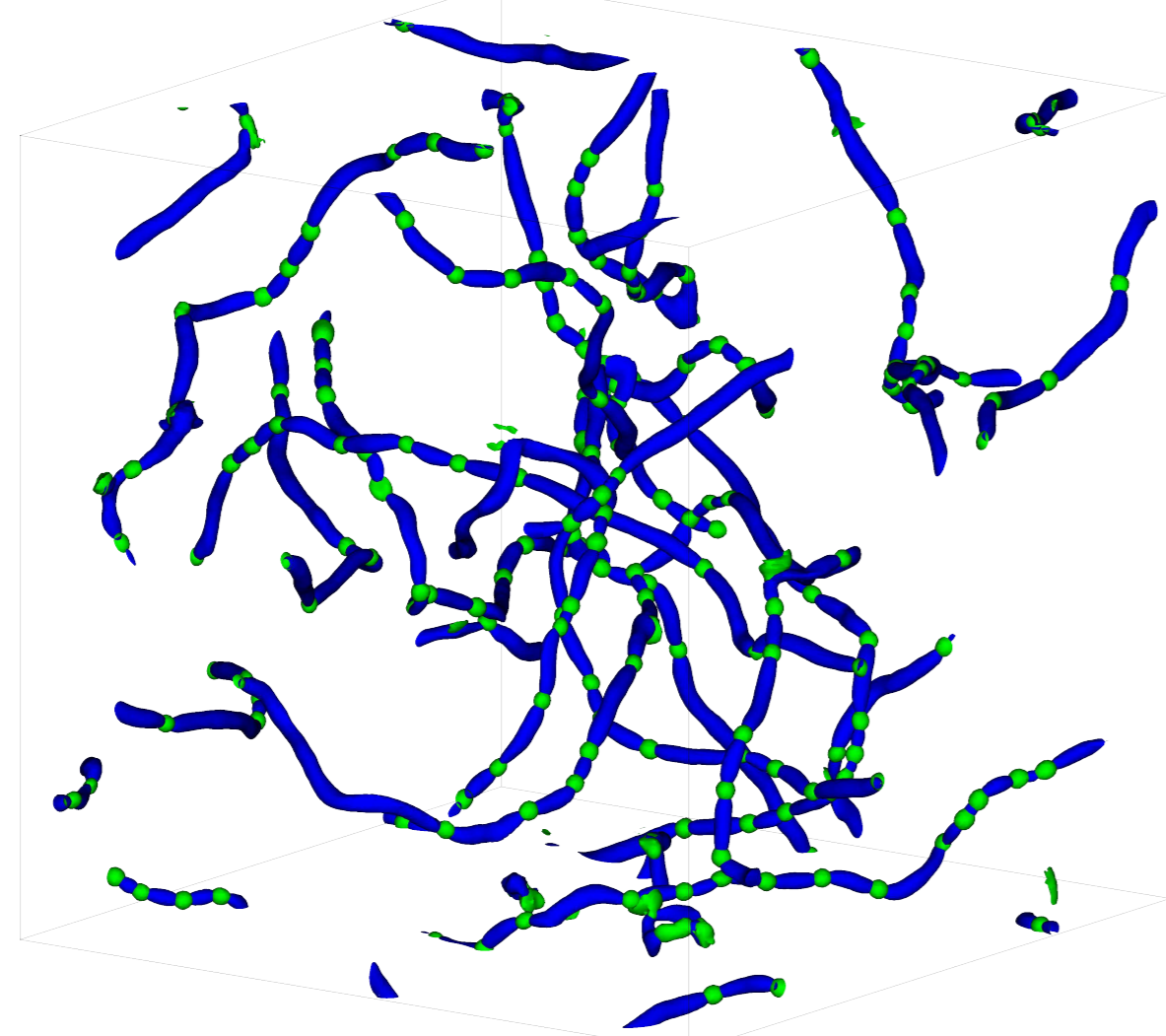
$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}^a F^{\mu\nu a} + \sum_n \text{Tr}[D_\mu \Phi_n][D^\mu \Phi_n] - V(\Phi_1, \Phi_2); \quad V(\Phi_1, \Phi_2) = -m_1^2 \text{Tr} \Phi_1^2 - m_2^2 \text{Tr} \Phi_2^2 + \lambda \left[(\text{Tr} \Phi_1^2)^2 + (\text{Tr} \Phi_2^2)^2 \right] + \kappa (\text{Tr} \Phi_1 \Phi_2)^2$$

where $D_\mu = \partial_\mu + igA_\mu$, $F_{\mu\nu} = F_{\mu\nu}^a \tau^a$ and $A_\mu = A_\mu^a \tau^a$, $\tau^a = \sigma^a/2$. Here, Φ_1 and Φ_2 are adjoint Higgs fields ($\Phi = \phi^a \sigma^a$).

- ▶ The system undergoes two symmetry breaking phase transitions, $SU(2) \rightarrow U(1) \rightarrow Z_2$. The first, $SU(2) \rightarrow U(1)$, creates 't Hooft-Polyakov monopoles with mass M , the second, $U(1) \rightarrow Z_2$, confines the flux to cosmic strings with tension μ , like beads on a wire.
- ▶ Simulations are performed in a comoving $V = 720^3$ box with lattice spacing $a = 1$, with Hubble damping corresponding to an expanding radiation-dominated universe.
- ▶ We determine the location of strings and monopoles within the box, yielding L , the total (Manhattan) length of string and N , the number of monopoles. These yield the average monopole and string separations $\xi_m = (V/N)^{1/3}$ and $\xi_s = (V/L)^{1/2}$. The monopole locations are also used to measure \bar{v}_m , root mean square monopole velocity.
- ▶ We compute monopole separation along the string $d = L/N$ and hence the ratio $r = M/\mu d$ measuring the importance of the monopoles for the string dynamics^{3,4,5}.

Case 1: $m_1^2 > m_2^2$

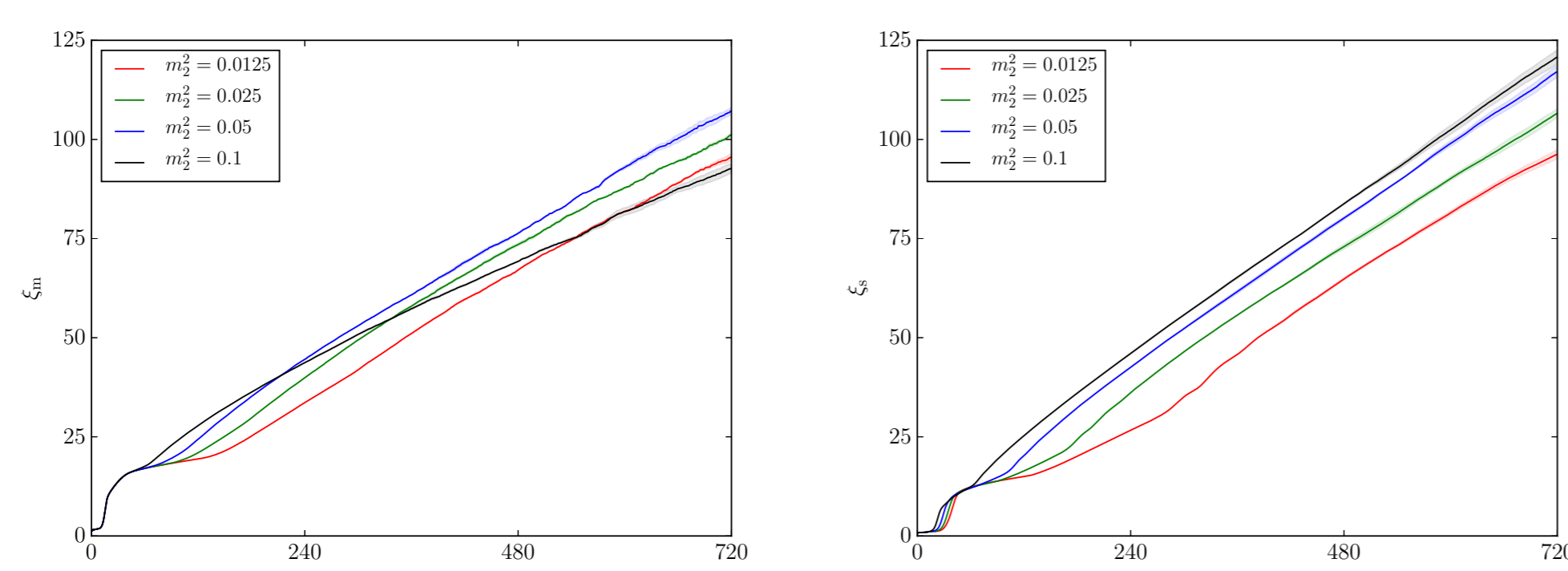
240^3 simulation with $m_1^2 = 0.25$, $m_2^2 = 0.1$, isosurfaces $\text{Tr} \Phi_1^2 = 0.2$, $\text{Tr} \Phi_2^2 = 0.04$, time $t = 240$.



- ▶ In this case, monopoles form as beads on the string.

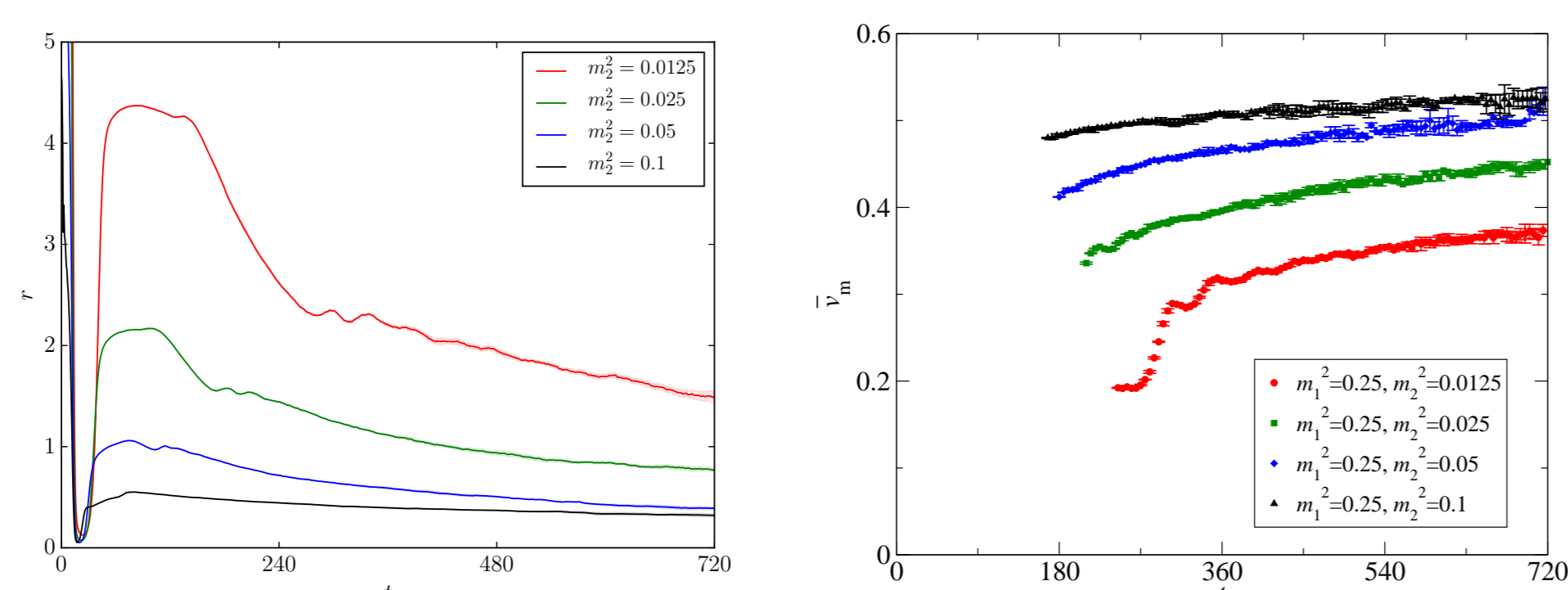


- ▶ We measured ξ_s and ξ_m and they grow linearly: a scaling network forms.



There is little apparent difference between the scaling for different μ .

- ▶ Our simulations show that r always decreases while \bar{v}_m increases, appearing to asymptote to a relativistic value.

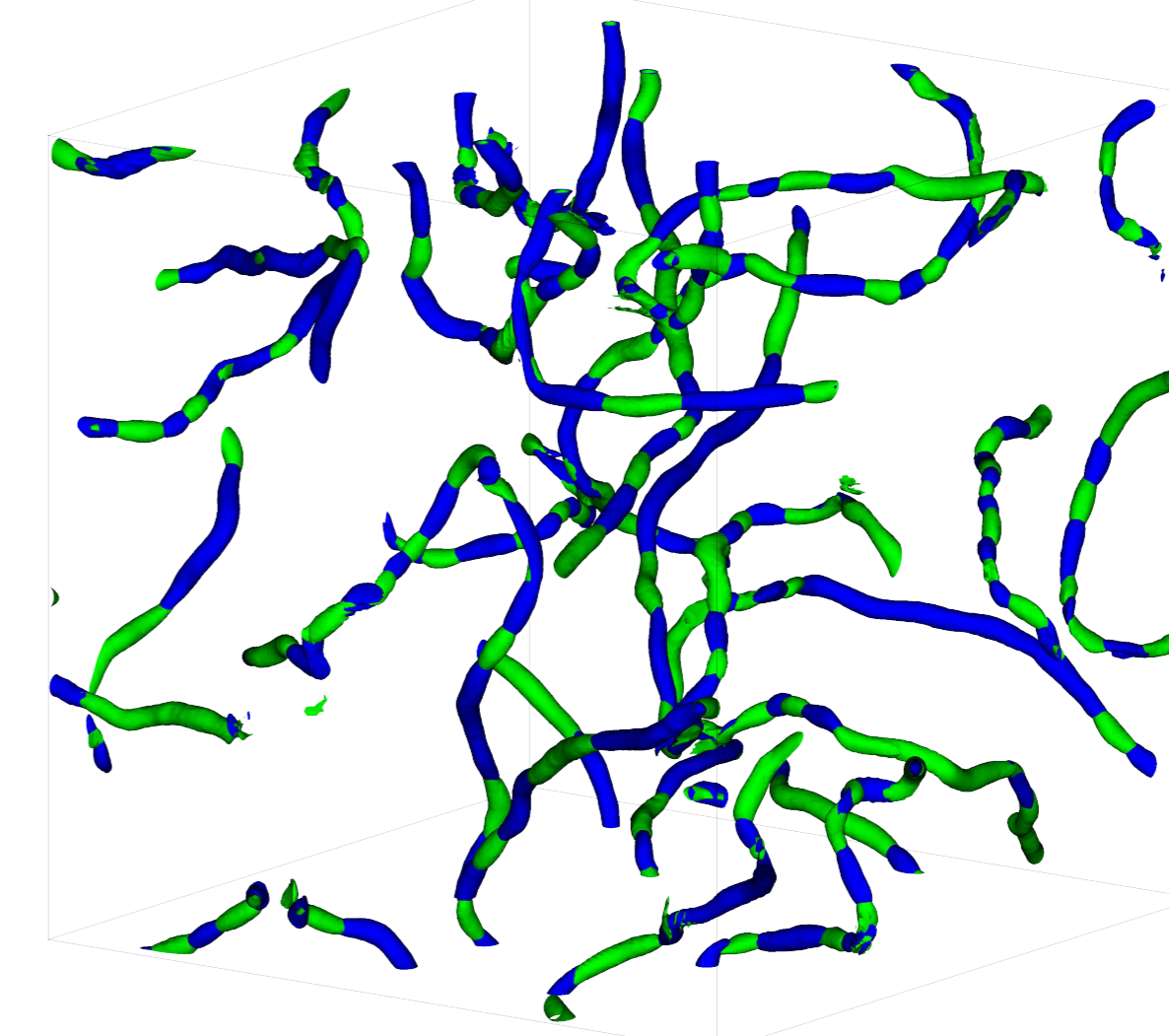


The monopole velocities are also in line with expected string velocities.

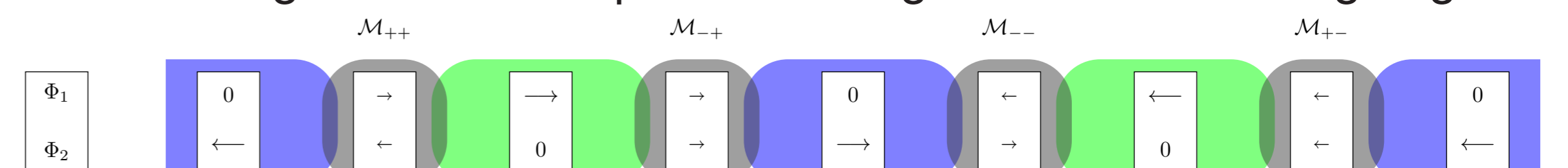
Key results: Scaling network forms; monopoles are unimportant; average monopole velocity does not decrease.

Case 2: $m_1^2 = m_2^2$

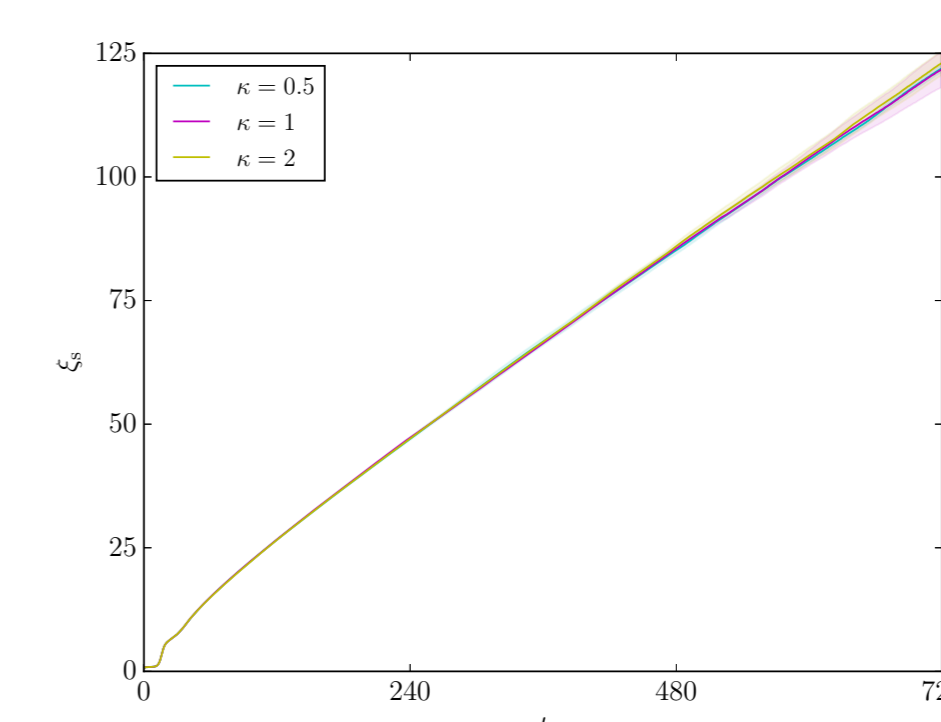
240^3 simulation with $m_1^2 = m_2^2 = 0.25$, $\kappa = 1$, isosurfaces $\text{Tr} \Phi_1^2 = \text{Tr} \Phi_2^2 = 0.2$, time $t = 240$.



- ▶ For all $\kappa > 0$ one gets 'half monopoles' forming at the end of string segments:



- ▶ There is a global symmetry between the two scalar fields, and the larger κ is, the less the two fields overlap.
- ▶ However, for all values of κ , the result is a scaling network of strings:



- ▶ We find that ξ_s scales with coefficient 0.16 ± 0.01 , which corresponds to string densities approximately 40% higher than in the abelian Higgs model. CMB constraints are therefore stronger for this model⁶.
- ▶ Because of the global symmetry, we cannot count the number of half monopoles.

Key results: When $m_1^2 = m_2^2$, we get novel strings with 'half monopole' structures. These still produce a scaling network.

References

1. M. Hindmarsh and T. Kibble, *Phys. Rev. Lett.* **55**, 2398 (1985)
2. T. Vachaspati and A. Vilenkin, *Phys. Rev. D* **35**, 1131 (1987)
3. V. Berezinsky and A. Vilenkin, *Phys. Rev. Lett.* **79**, 5202 (1997)
4. J. J. Blanco-Pillado and K. D. Olum, *JCAP* **1005**, 014 (2010)
5. C. J. A. P. Martins, *Phys. Rev. D* **82**, 067301 (2010)
6. N. Bevis *et al.*, *Phys. Rev. D* **82**, 065004 (2010)